

A Model of Stealth Dark Matter



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Based on 1402.6656, 1503.04203, 1503.04205
with LSD Collaboration

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Lattice Strong Dynamics Collaboration

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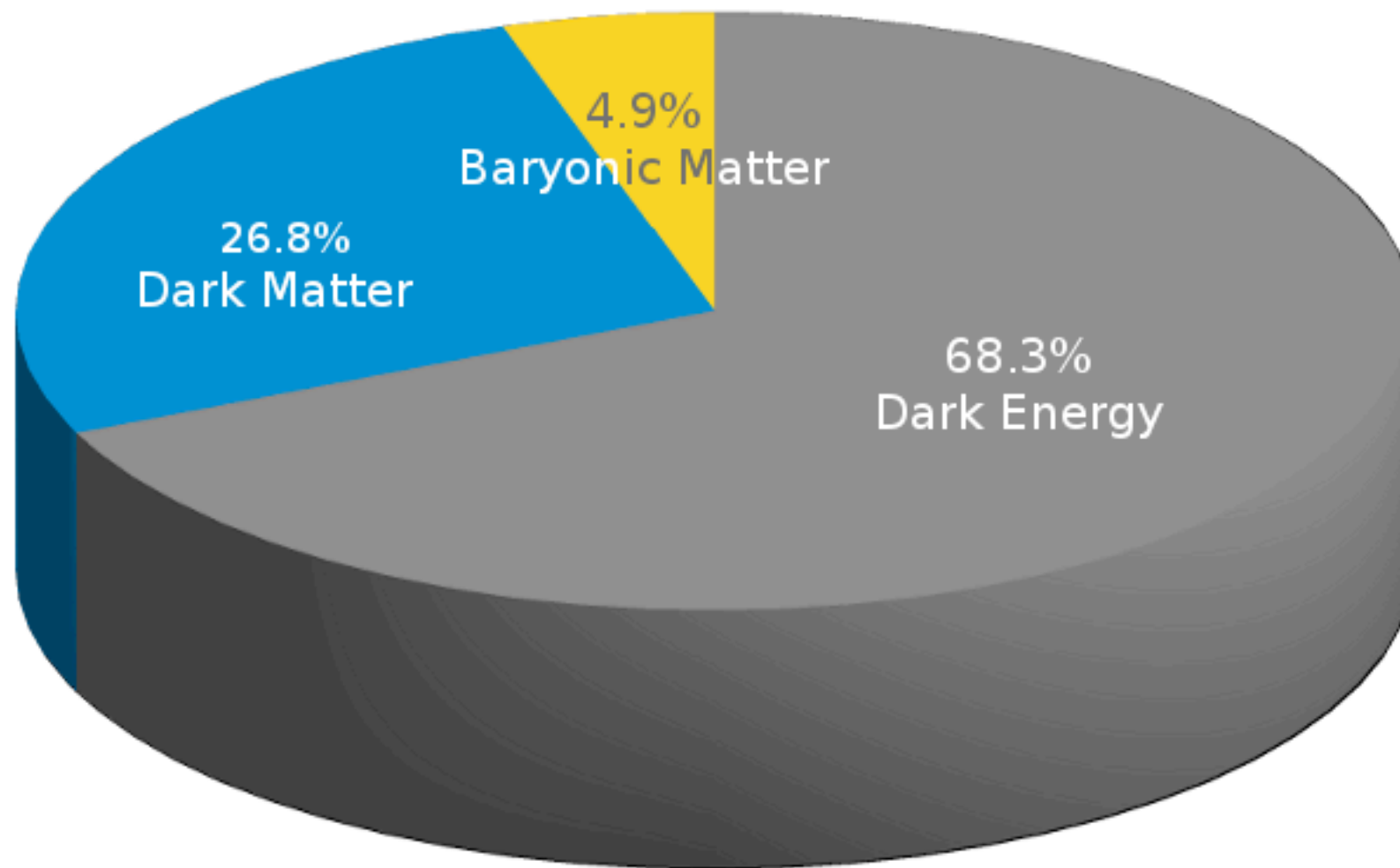
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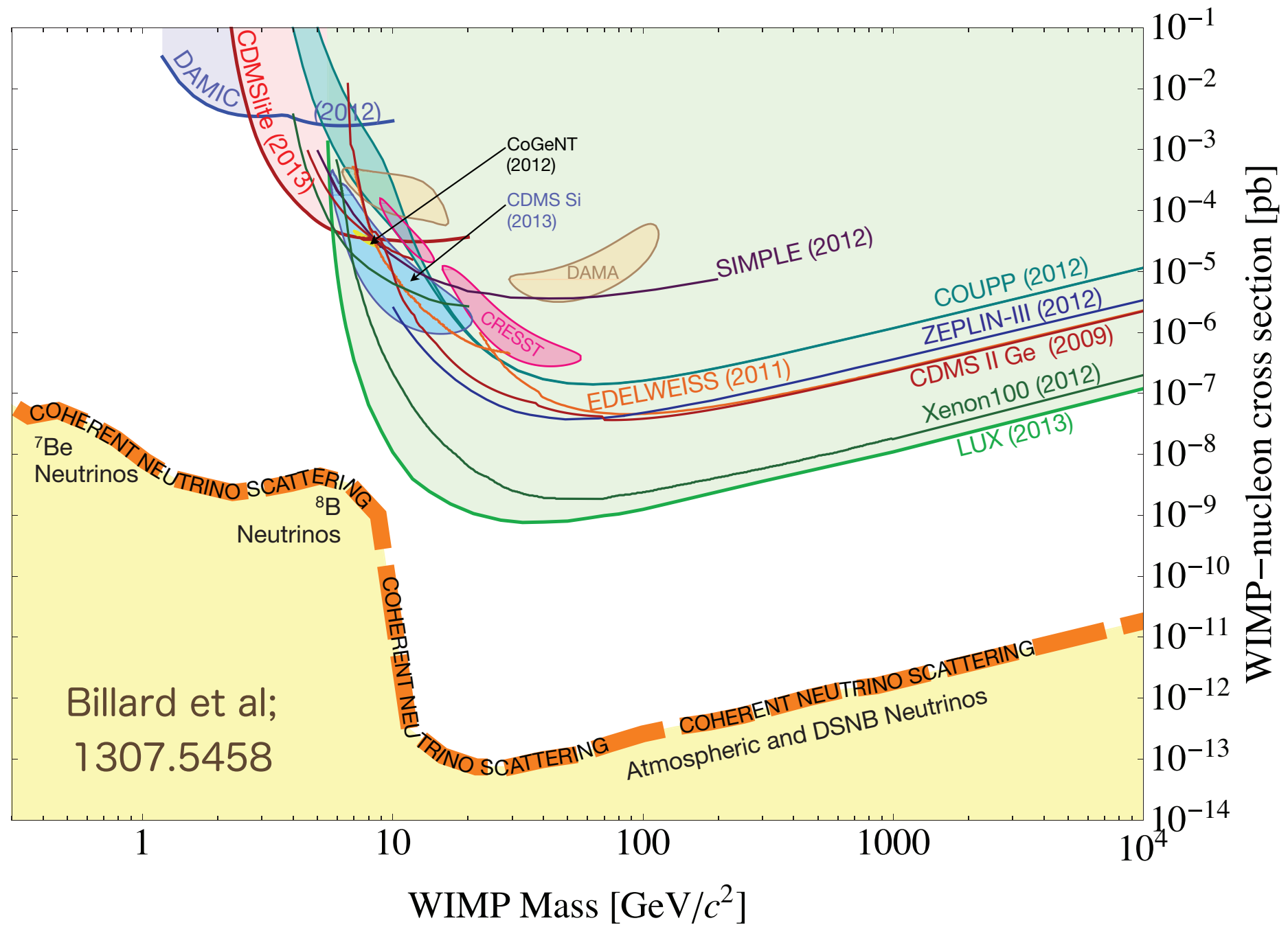
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Dark Matter



Overwhelming evidence from CMB; galaxies; clusters; BAO; ...

Direct Detection Cross Section



Suppressed Cross Sections

σ_{nucleon}

σ_{RADAR}

4th Dirac neutrino

B-52



$$\sim 10^{-38} \text{ cm}^2$$

$$\sim 100 \text{ m}^2$$

Quirky DM

Falcon

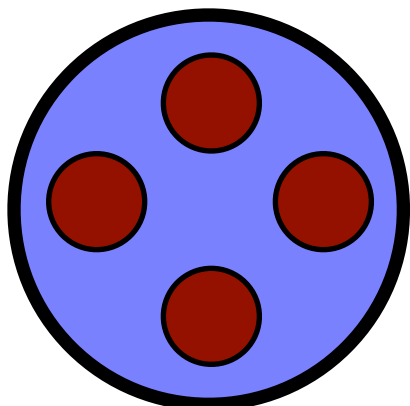


$$\sim 10^{-43} \text{ cm}^2$$

$$\sim 10^{-2} \text{ m}^2$$

Stealth DM

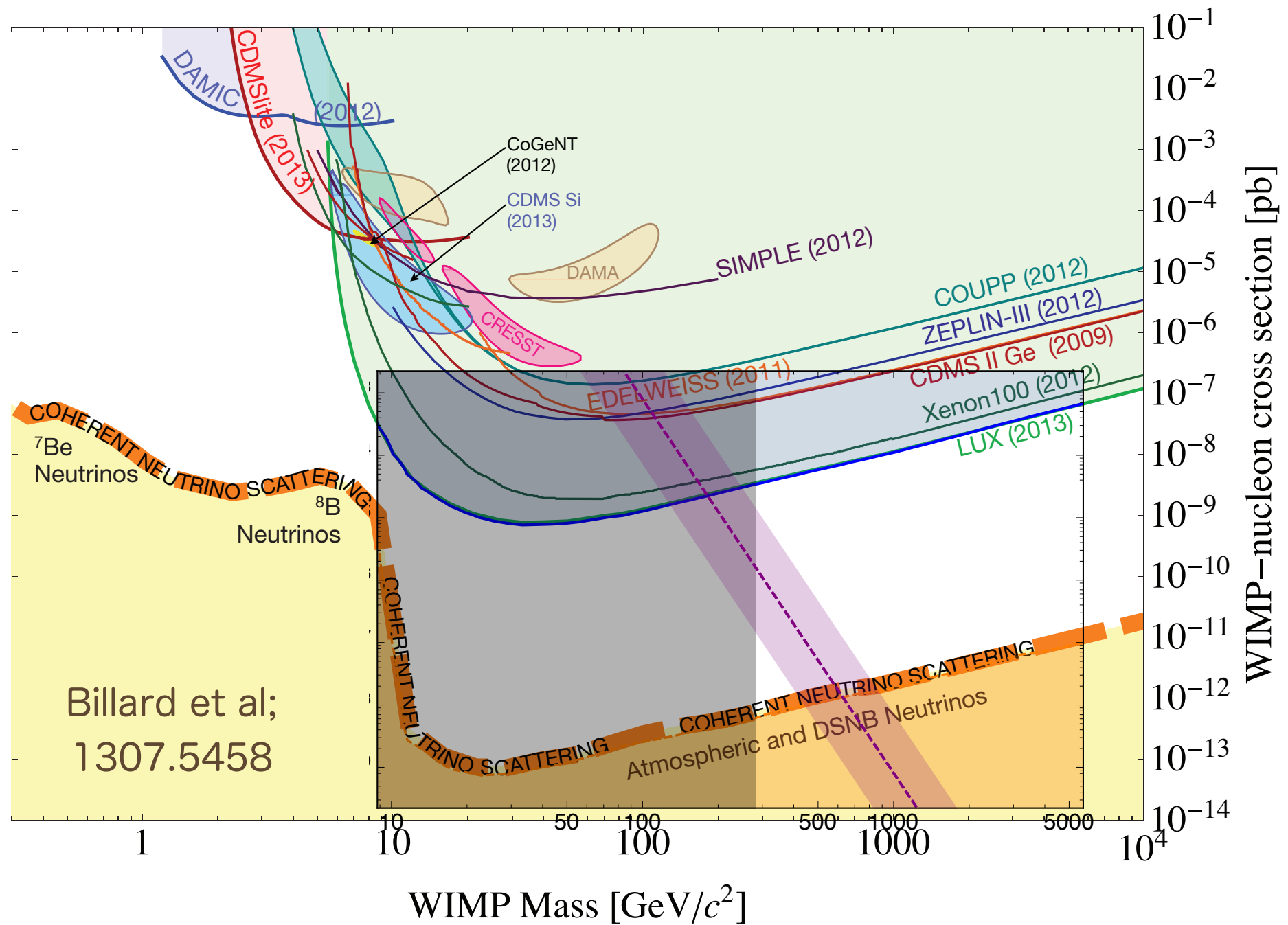
Stealth B-2



$$\sim \left(\frac{200 \text{ GeV}}{m_B} \right)^6 \times 10^{-45} \text{ cm}^2$$

even smaller

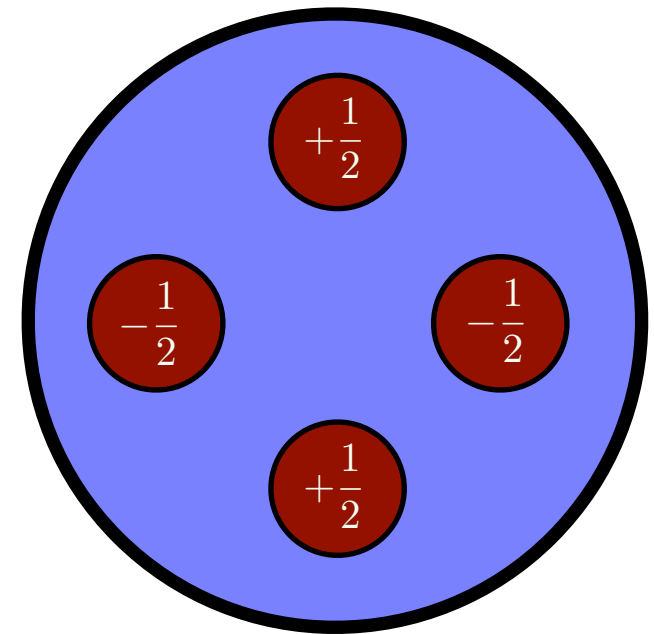
Direct Detection Cross Section



Effective lower bound on composite DM with electrically charged constituents.

Stealth Dark Matter

- Scalar baryon of strongly-coupled $SU(N_D)$, with N_D even [focus on $SU(4)$] and dark fermions transforming under EW group
- All mass scales are technically natural;
very roughly $100 \text{ GeV} \lesssim M_f \sim \Lambda_D \lesssim 100 \text{ TeV}$
- We use lattice simulations to calculate several non-perturbative observables (mass spectrum; interactions of DM with SM)
- Naturally “stealthy” with respect to direct detection; we determine the “ultimate” lower bound on composite DM with charged constituents
- LHC phenomenology completely different from weakly-coupled DM models



Lattice Gauge Theory Simulations

Ideal tool to calculate properties of theories with

$$M_f \sim \Lambda_D$$

in the fully non-perturbative regime. Joy of these calculations is that what we simulate **is** interesting “out of the box” without chiral extrapolations.

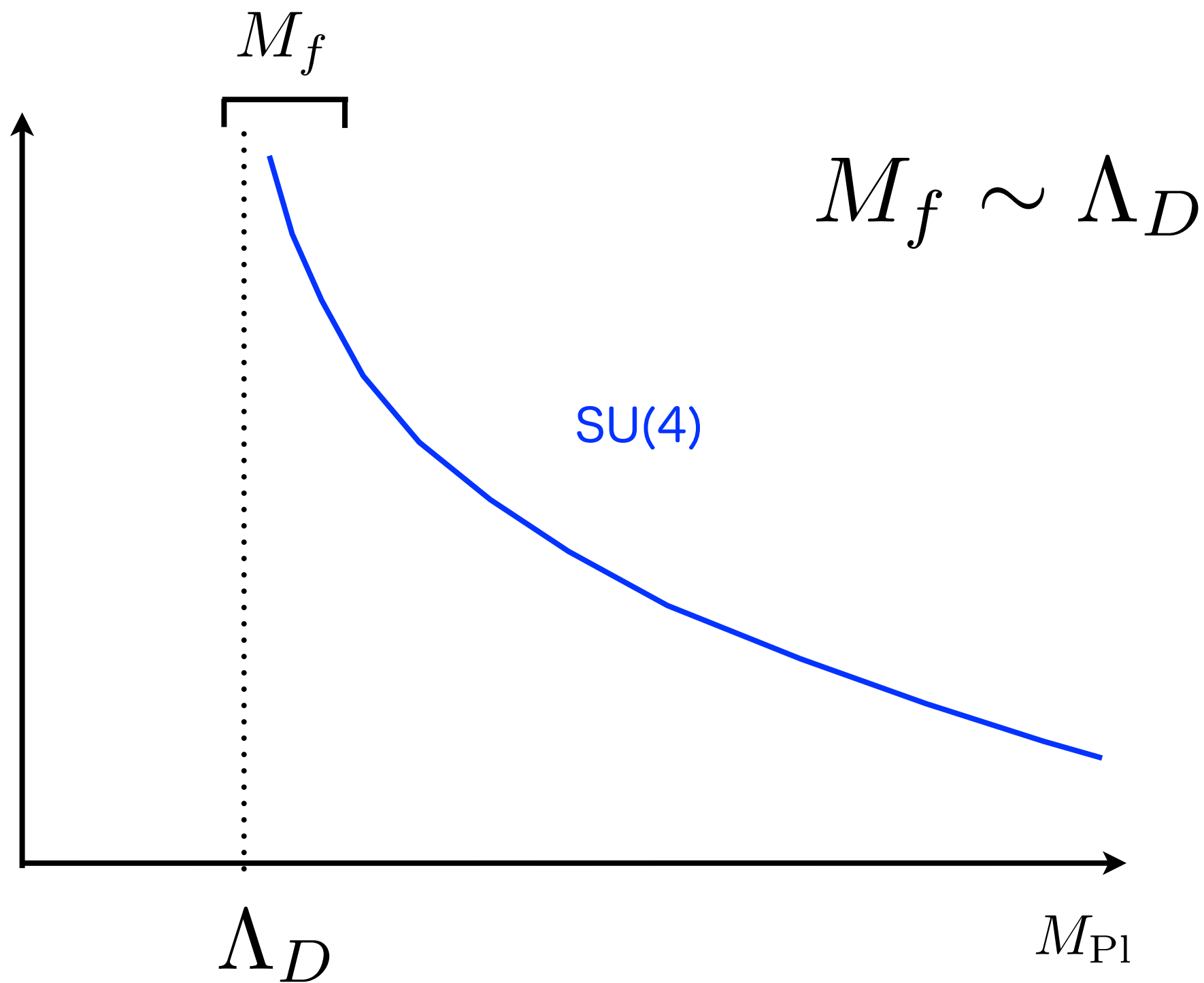


Relevant to DM: Thus far, we have accurate estimates of the spectrum, the “sigma term”, and polarizability. Future work will nail down additional correlators (for S parameter), meson form factor, ...

Simulated with modified Chroma mainly on LLNL sequoia/vulcan. Quenched, unmodified Wilson fermions. Several volumes and lattice spacings.

Dynamics

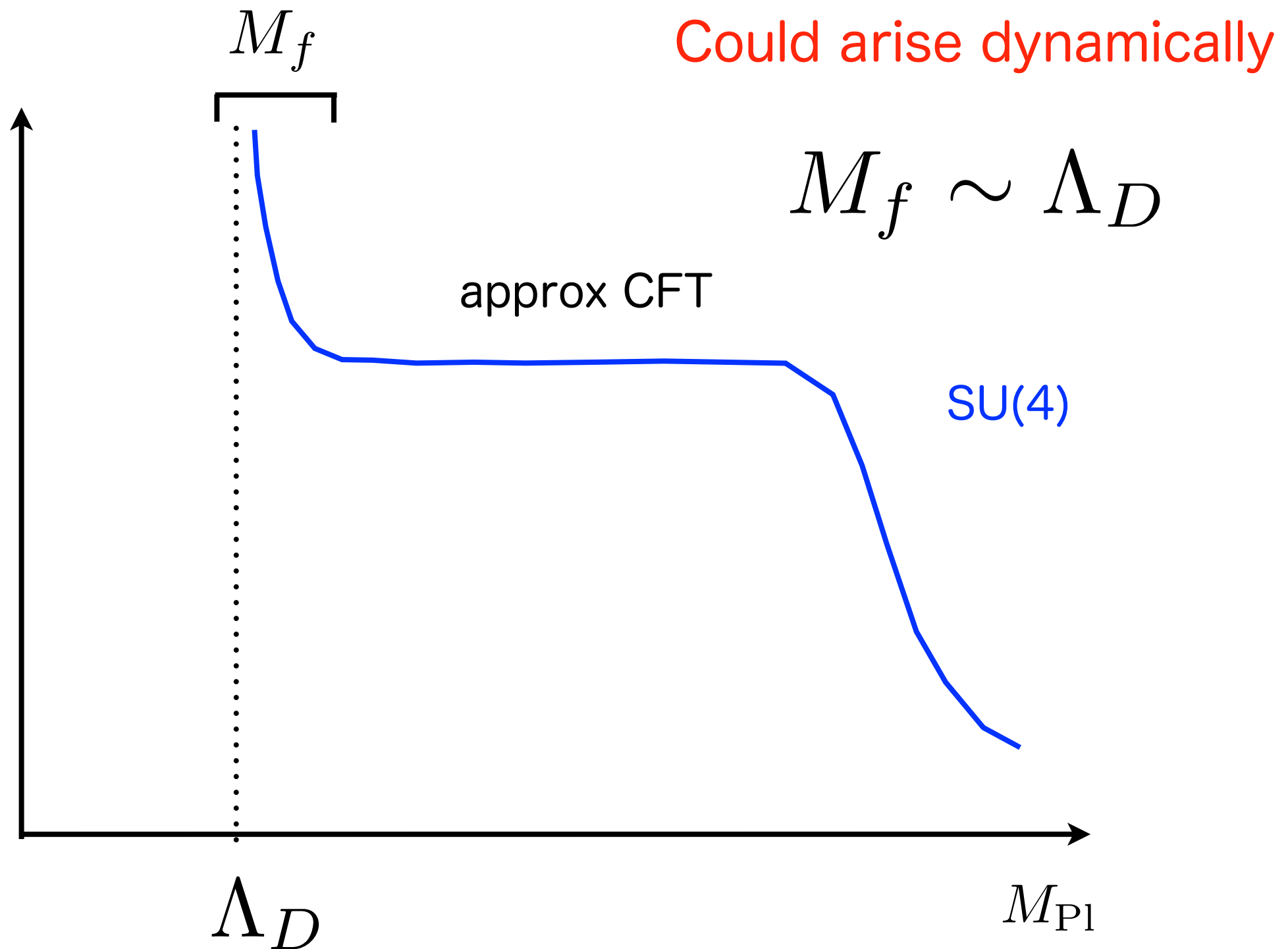
Dark fermions



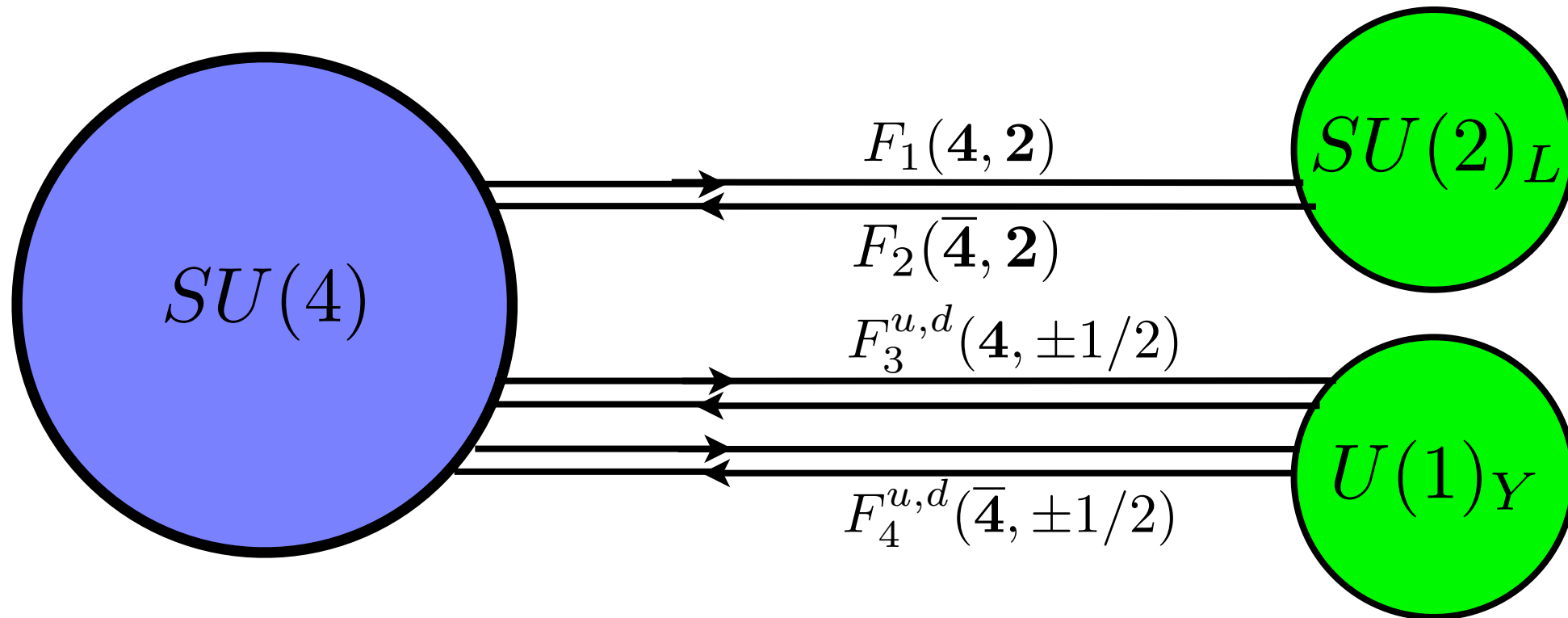
Dynamics

Dark fermions

Could arise dynamically



Dark Fermions



Vector-like masses

$$M_{12}\epsilon_{ij}F_1^iF_2^j - M_{34}^uF_3^uF_4^d + M_{34}^dF_3^dF_4^u + h.c.,$$

EW breaking masses

$$\begin{aligned} & y_{14}^u\epsilon_{ij}F_1^iH^jF_4^d + y_{14}^dF_1 \cdot H^\dagger F_4^u \\ & - y_{23}^d\epsilon_{ij}F_2^iH^jF_3^d - y_{23}^uF_2 \cdot H^\dagger F_3^u \end{aligned} + h.c.$$

Dark Flavor Symmetries

Under SU(4): $U(4) \times U(4)$

Weak gauging: $[SU(2) \times U(1)]^4$ (that contains $SU(2)_L \times U(1)_Y$)

Vector-like masses: $SU(2)_L \times U(1)_Y \times U(1) \times U(1)$

Yukawas with Higgs: $U(1)_B$

Dark baryon number automatic.

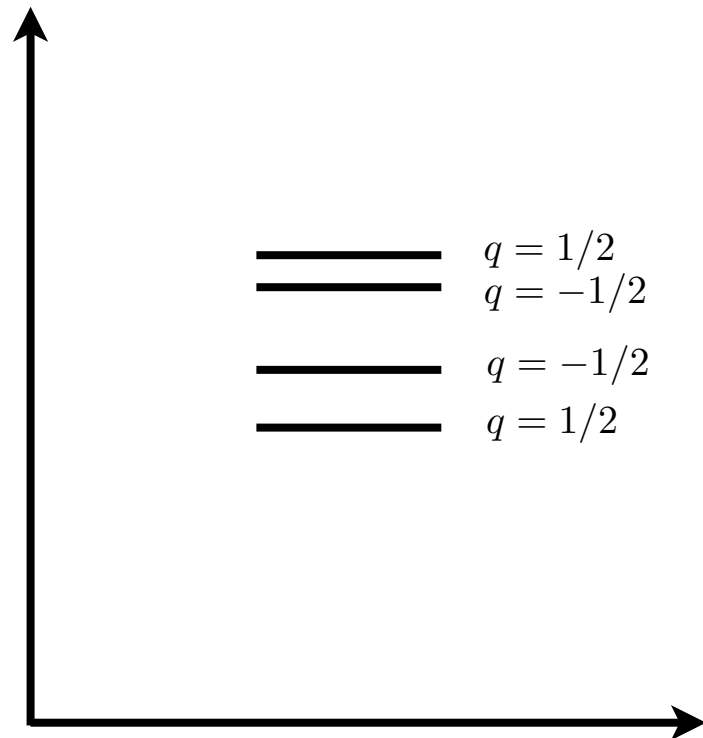
and **very safe** against cutoff scale violations of global symmetries
e.g.

$$\frac{qqqq H^\dagger H}{\Lambda_{\text{cutoff}}^4}$$

[This is one reason to prefer SU(4) over SU(2).]

Dark Fermion Mass Spectrum

General

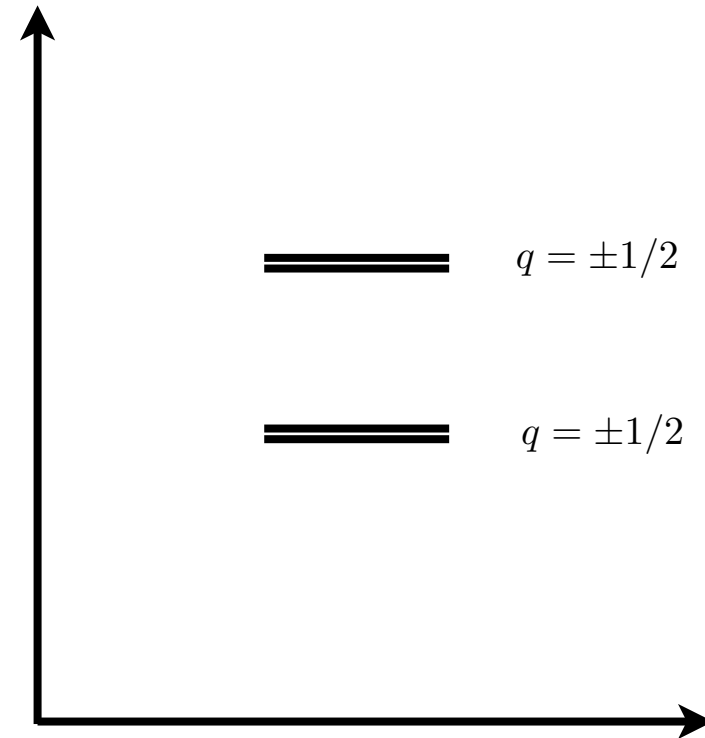


$$M_{12}, M_{34}^u, M_{34}^d$$

$$y_{14}^u, y_{14}^d$$

$$y_{23}^u, y_{23}^d$$

Custodial SU(2)



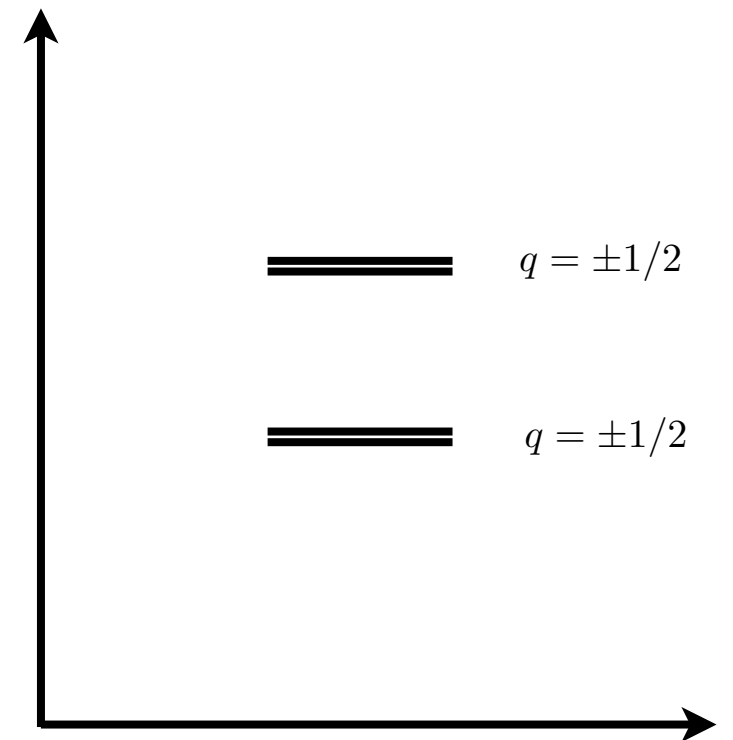
$$M_{34}^u = M_{34}^d$$

$$y_{14}^u = y_{14}^d$$

$$y_{23}^u = y_{23}^d$$

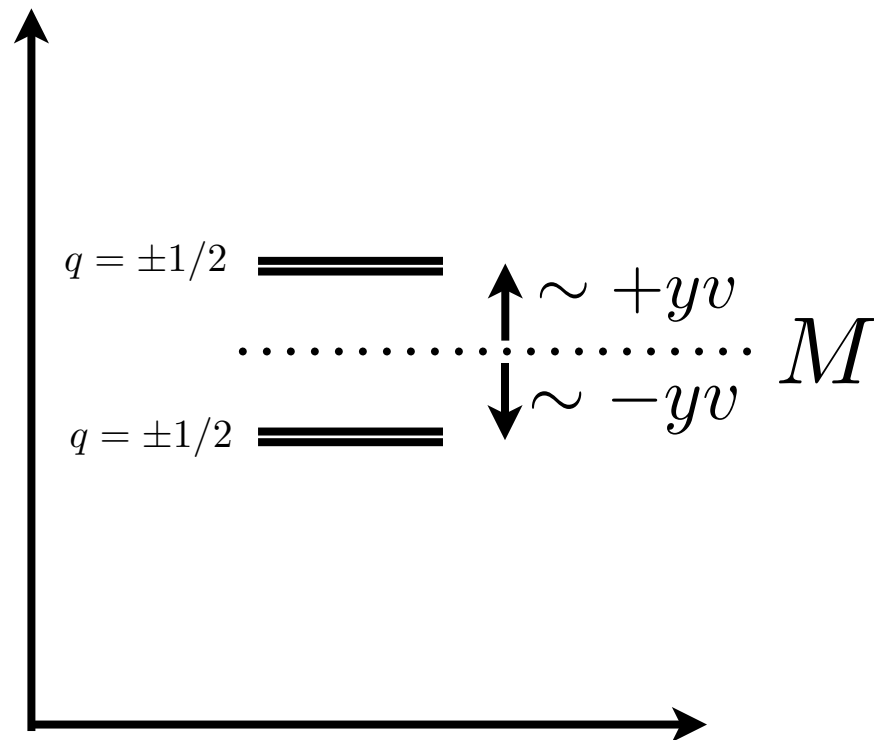
Custodial SU(2)

- Lightest baryon is a **neutral complex scalar**
(eliminates operators dependent on spin,
e.g., dim-5 magnetic moment)
- Contributions to **T parameter vanish**
(no need to make life more complicated)
- Weak isospin exactly **zero**
(no Z coupling to dark matter; otherwise significant constraints)
- Dim-6 charge radius **vanishes**
(more stealthy w.r.t. direct detection;
one less thing to calculate on lattice)

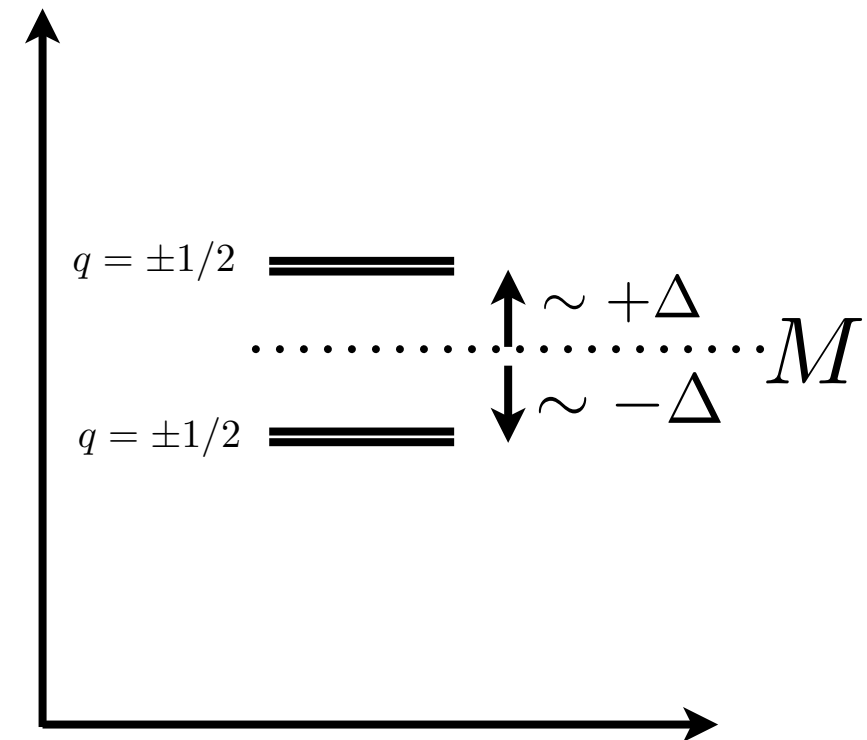


Two Distinct “Cases”

“Linear Case”



“Quadratic Case”



As we’ll see, Higgs boson coupling to lightest dark fermions is proportional to

y

Linear Case

y^2

Quadratic Case

A similar observation of linear/quadratic effect also in Hill, Solon; 1401.3339

Approximately Symmetric / Vector-Like

Fermion mass matrices with custodial SU(2)

$$M^u = M^d = \begin{pmatrix} M \pm \Delta & y_{14}v/\sqrt{2} \\ y_{23}v/\sqrt{2} & M \mp \Delta \end{pmatrix}$$

Convenient to expand around the symmetric matrix limit

$$y_{14} = y + \epsilon_y$$

$$y_{23} = y - \epsilon_y$$

Then the axial current

$$j_{+,\text{axial}}^\mu \supset c_{\text{axial}} \overline{\Psi}_1^u \gamma^\mu \gamma_5 \Psi_1^d$$

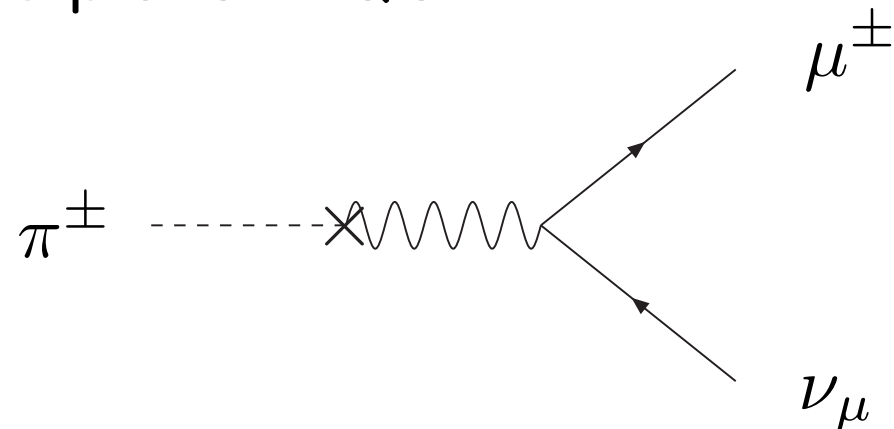
becomes

$$c_{\text{axial}} = \frac{\epsilon_y y v^2}{2M \sqrt{2\Delta^2 + y^2 v^2}}$$

$$\simeq \frac{\epsilon_y v}{2M} \times \begin{cases} 1 & \text{Linear Case} \\ yv/(\sqrt{2}\Delta) & \text{Quadratic Case.} \end{cases}$$

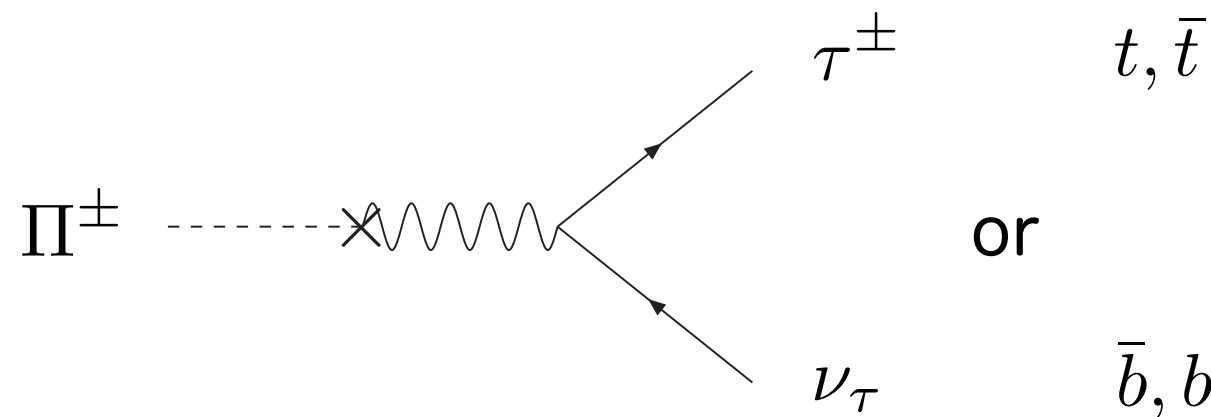
Charged Meson Decay

Like pions in QCD



$$\langle 0 | j_{\pm, \text{axial}}^{\mu} | \pi^{\pm} \rangle = i f_{\pi} p^{\mu}$$

Lightest dark mesons **decay** through



$$\langle 0 | j_{\pm, \text{axial}}^{\mu} | \Pi^{\pm} \rangle = i f_{\Pi} p^{\mu}$$

The non-zero Yukawa couplings with $\epsilon_y \neq 0$ cause $j_{\pm, \text{axial}}^{\mu} \neq 0$

$$\frac{\Gamma(\Pi^+ \rightarrow f \bar{f}')}{\Gamma(\pi \rightarrow \mu^+ \nu_{\mu})} \simeq \frac{c_{\text{axial}}^2}{|V_{ud}|^2} \left(\frac{f_{\Pi}}{f_{\pi}} \right)^2 \left(\frac{m_f}{m_{\mu}} \right)^2 \left(\frac{m_{\Pi}}{m_{\pi}} \right)$$

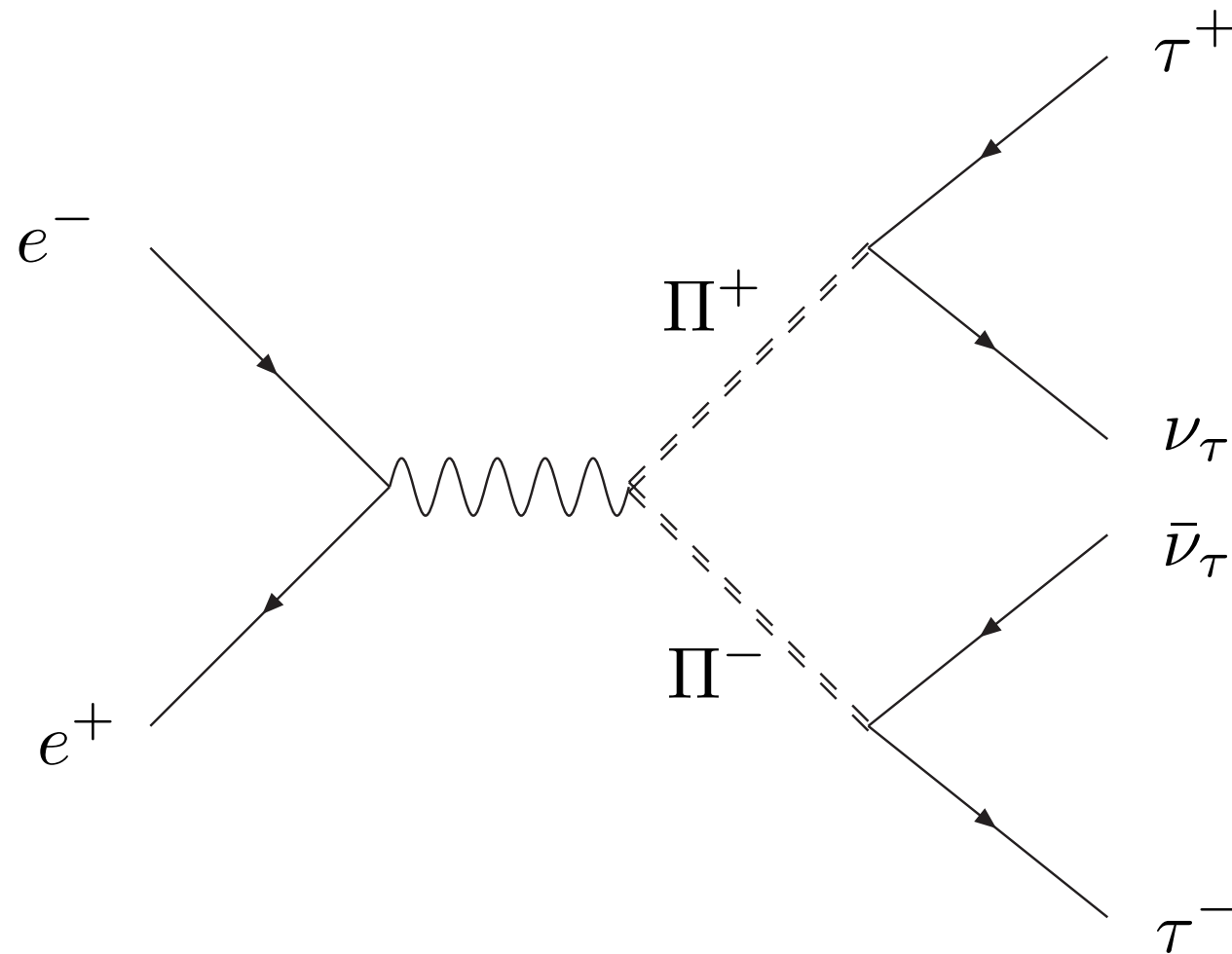
(unlike “Vector-like Confinement”)

Kilic, Okui, Sundrum; 0906.0577

and so **dark mesons decay much faster** than QCD pions even with $c_{\text{axial}} \ll 1$

Lower bound on meson mass ...

Charged pion production at LEP II

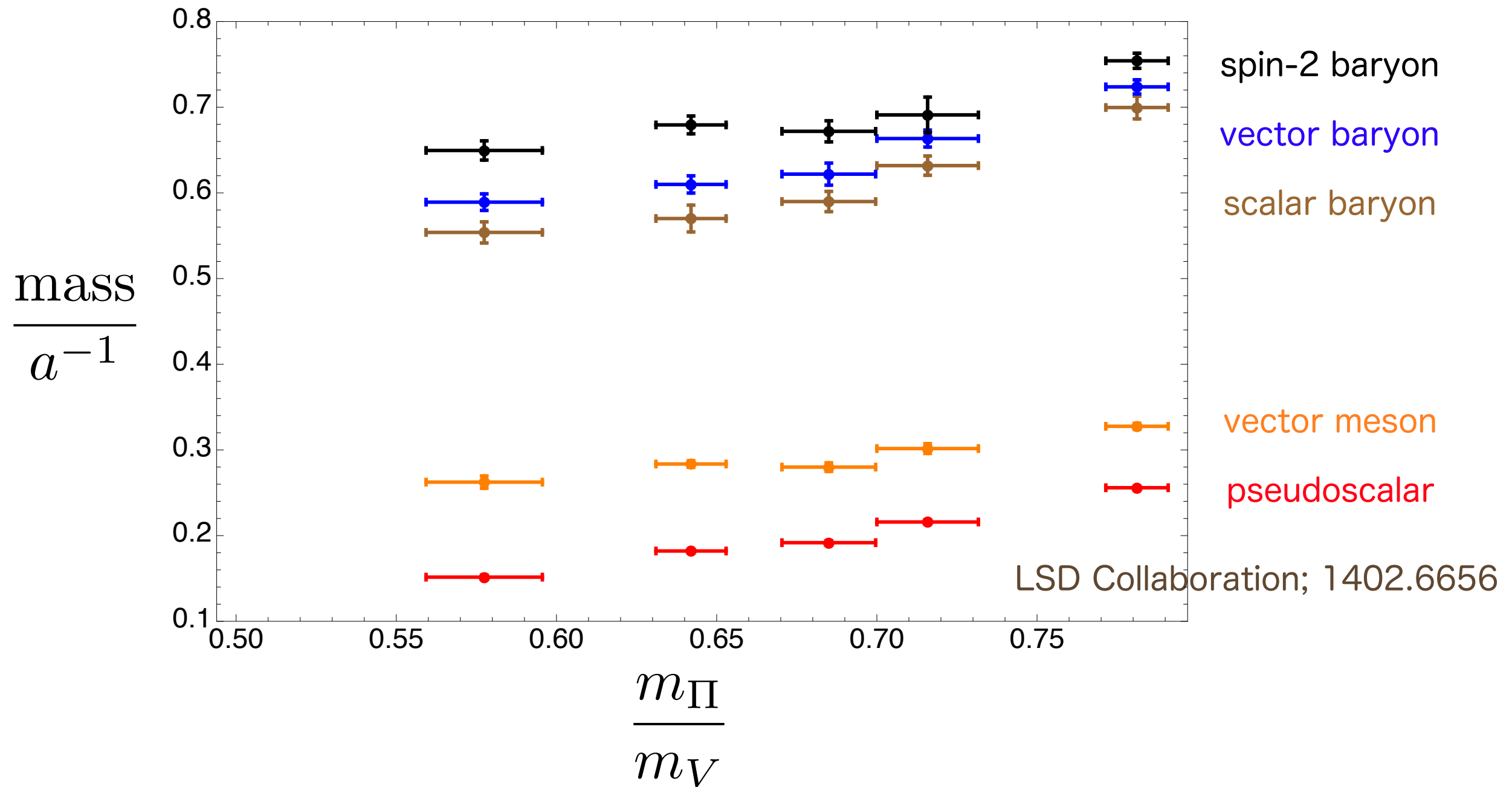


Using a crude recasting of bounds on staus, we find

$$m_{\Pi^\pm} > 86 \text{ GeV}$$

This is fairly robust to promptness/non-promptness of dark meson decay.

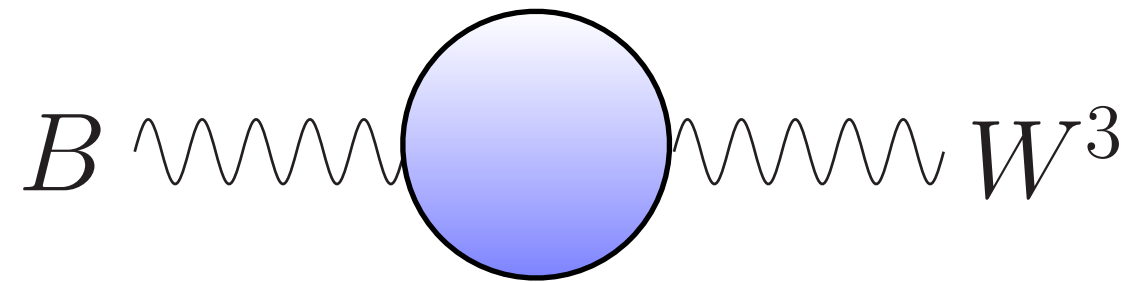
... becomes lower bound on the baryon mass



Within the range simulated on our lattices, we obtain

$$2.5 \lesssim \frac{m_B}{m_{\Pi}} \lesssim 3.8$$

S parameter



Peskin, Takeuchi (1990, 92)

Obviously $\Delta S \rightarrow 0$ as $(yv) \rightarrow 0$.

With custodial SU(2), approximate symmetric, and M_1 close to M_2

$$S \propto \int d^4x e^{-i\mathbf{q}\cdot\mathbf{x}} \langle j_3^\mu(x) j_Y^\nu(0) \rangle \simeq \frac{\epsilon_y^2 v^2}{4M^2} G_{LR}^{\mu\nu},$$

\uparrow
 $G_{LR}^{\mu\nu} \equiv \langle \bar{\psi}^u \gamma^\mu P_L \psi^u \bar{\psi}^u \gamma^\nu P_R \psi^u \rangle|_{\text{connected}}$

and thus can be **easily** suppressed below experimental limits.

[Vector-like masses for dark fermions **crucial**.]

Effective Higgs Coupling

The Higgs coupling to the lightest dark fermions

$$\mathcal{L} \supset y_\Psi h \bar{\Psi}_1 \Psi_1$$

$$y_\Psi = \frac{y^2 v}{M_2 - M_1} + O(\epsilon_y) \simeq \begin{cases} \frac{y}{\sqrt{2}} & \text{Linear Case} \\ \frac{y^2 v}{2\Delta} & \text{Quadratic Case.} \end{cases}$$

This leads to an **effective Higgs coupling** to the dark scalar baryon

$$g_B \simeq f_f^B \times \begin{cases} y_{\text{eff}} & \text{Linear Case} \\ y_{\text{eff}}^2 \frac{v}{m_B} & \text{Quadratic Case} \end{cases}$$

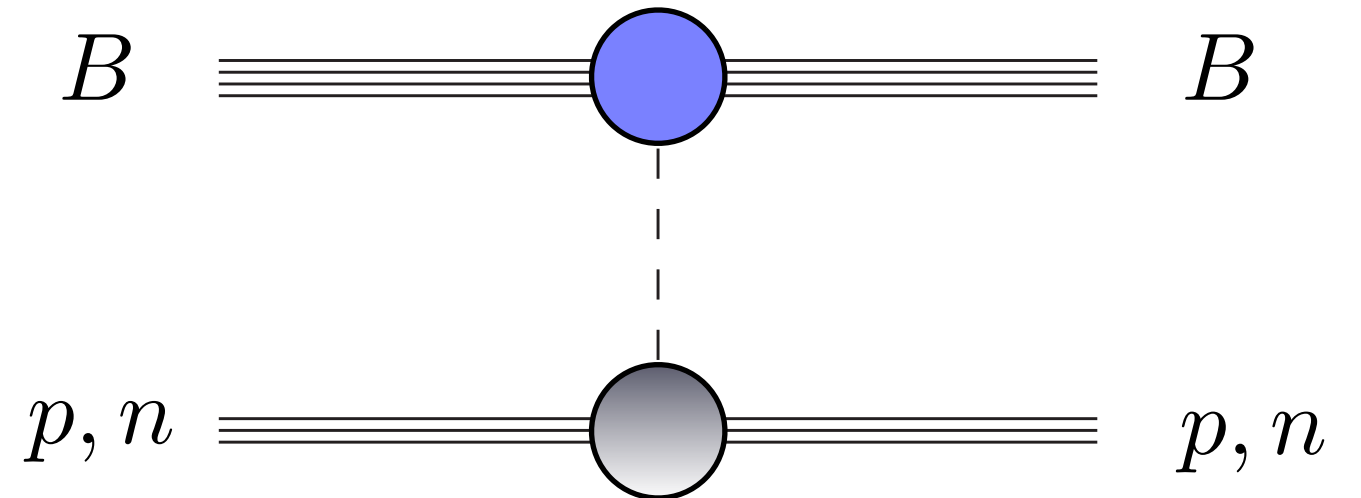
$$y_{\text{eff}} \equiv \begin{cases} y \frac{m_B}{\sqrt{2} M_1} & \text{Linear Case} \\ y \frac{m_B}{\sqrt{2} \Delta M_1} & \text{Quadratic Case.} \end{cases}$$

$$\langle B | m_f \bar{f} f | B \rangle = m_B f_f^B$$

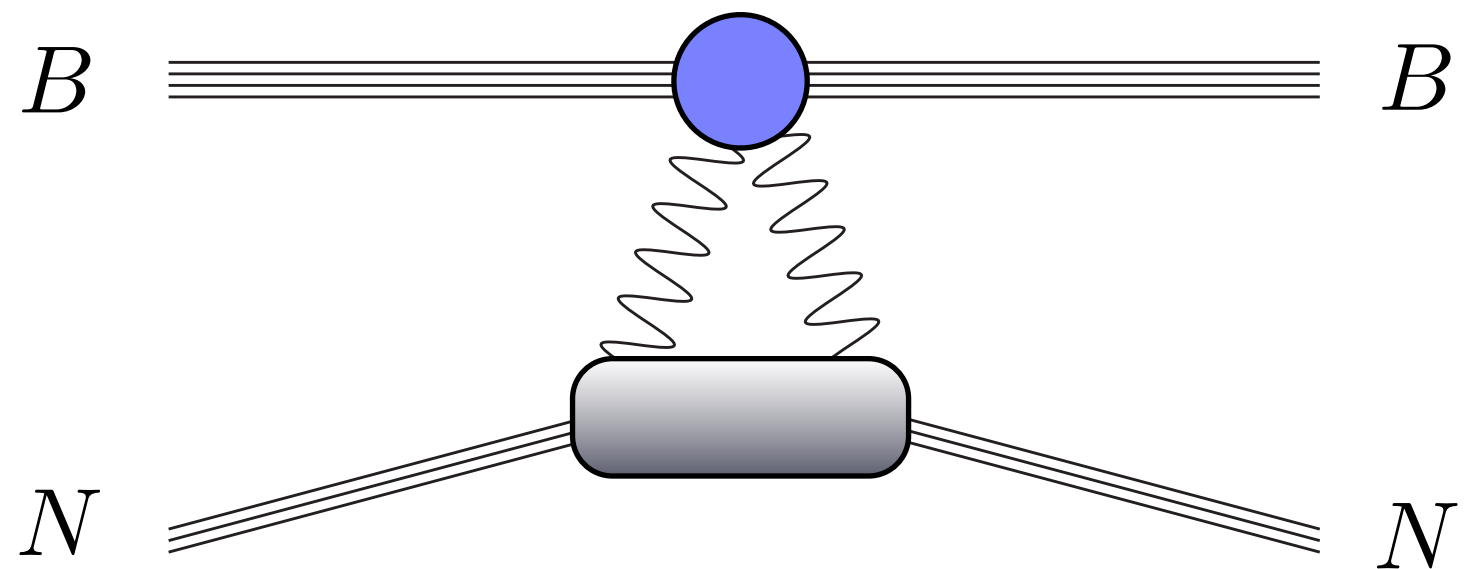
Extracted from **lattice!**

Direct Detection

Higgs exchange



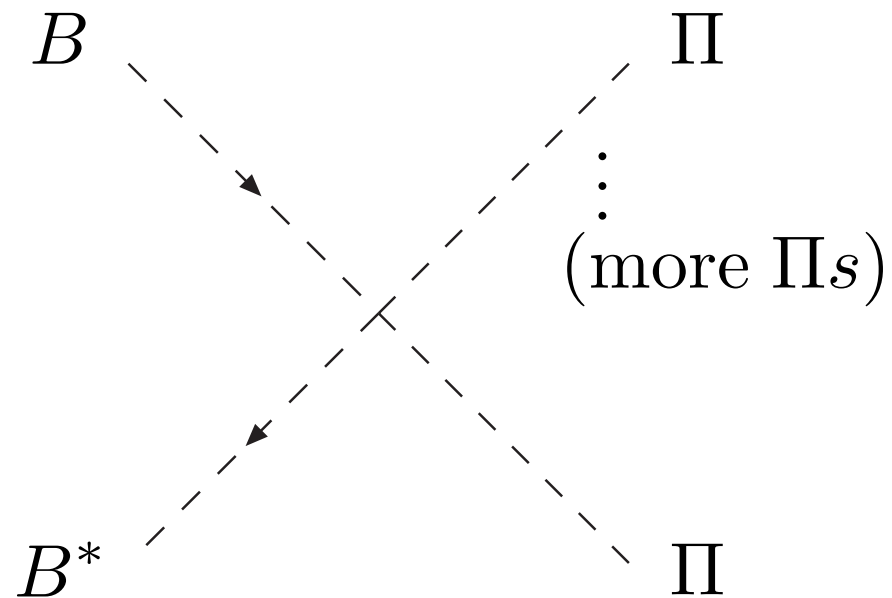
EM polarizability



Stay tuned for Enrico's talk!

Abundance

Symmetric



If $2 \rightarrow 2$ dominates the thermal annihilation rate and saturates unitarity, expect

Griest, Kamionkowski; 1990

$$m_B \sim 100 \text{ TeV}$$

Unfortunately, this is a **hard** calculation to do using lattice...

Asymmetric

e.g., through EW sphalerons

Barr, Chivukula, Farhi; 1990

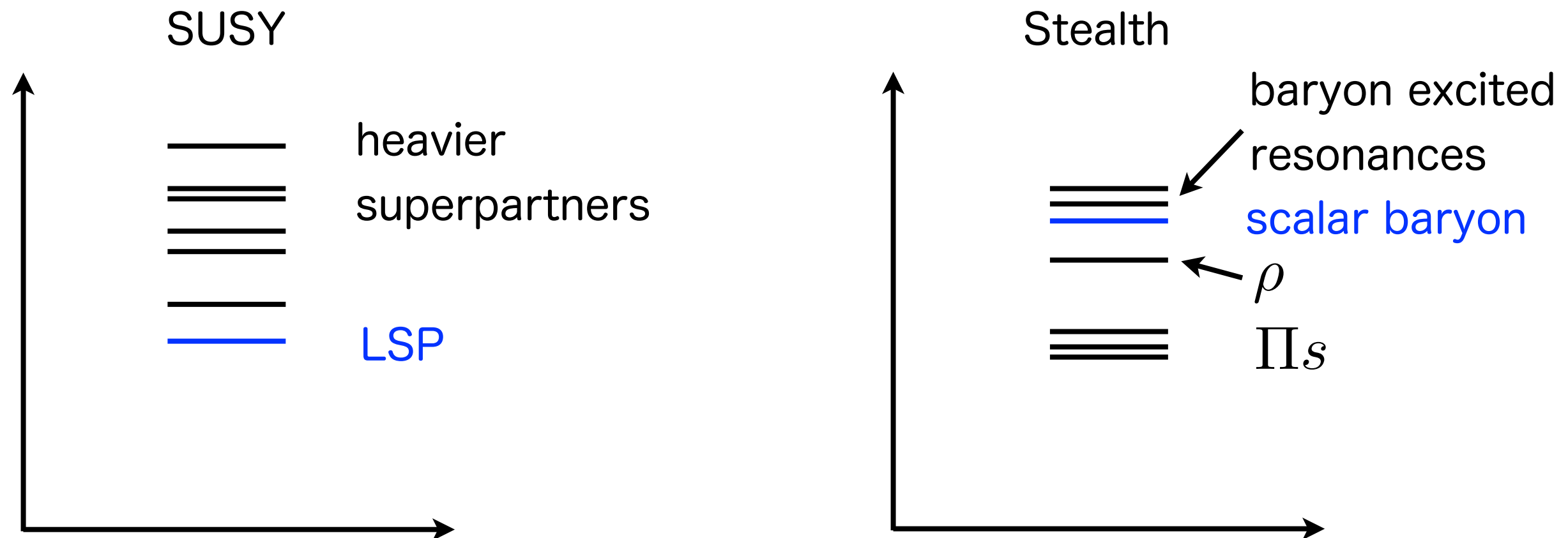
$$n_D \sim n_B \left(\frac{yv}{m_B} \right)^2 \exp \left[-\frac{m_B}{T_{\text{sph}}} \right]$$

IF EW breaking comparable to EW preserving masses, expect roughly

$$m_B \lesssim m_{\text{techni-B}} \sim 1 \text{ TeV}$$

How much less depends on several factors...

Colliders

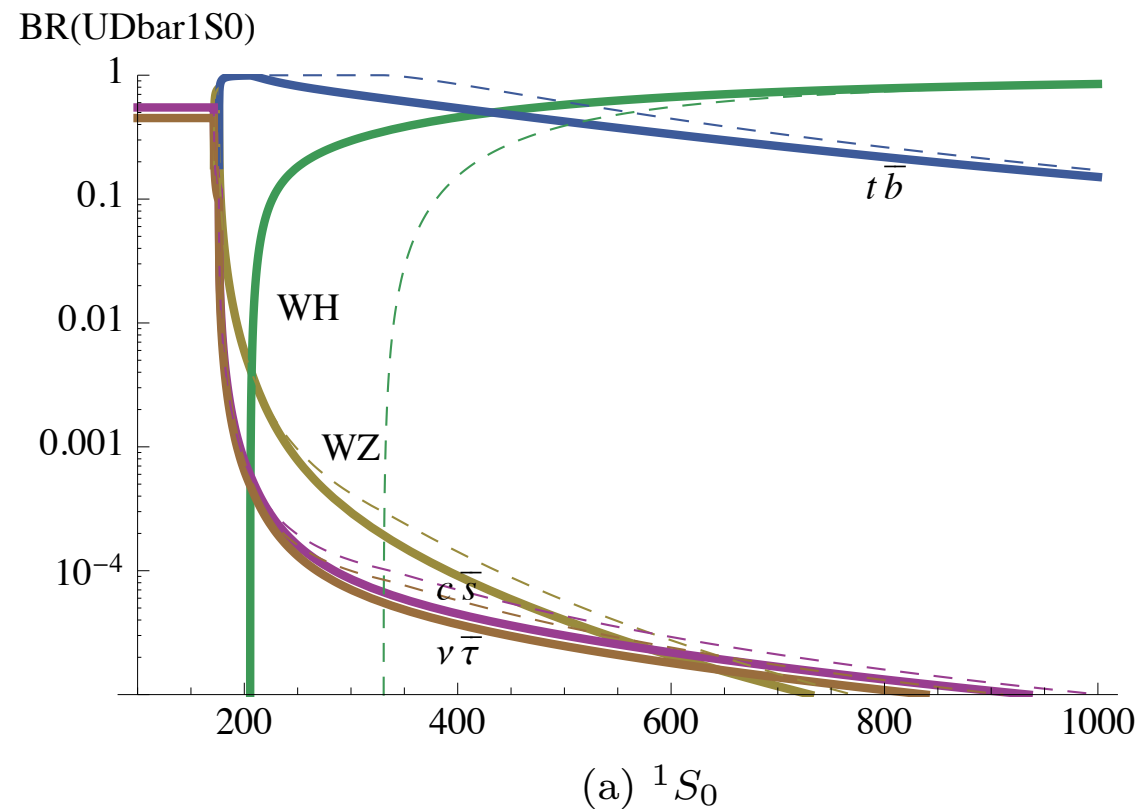


Collider searches dominated by light meson production and decay.

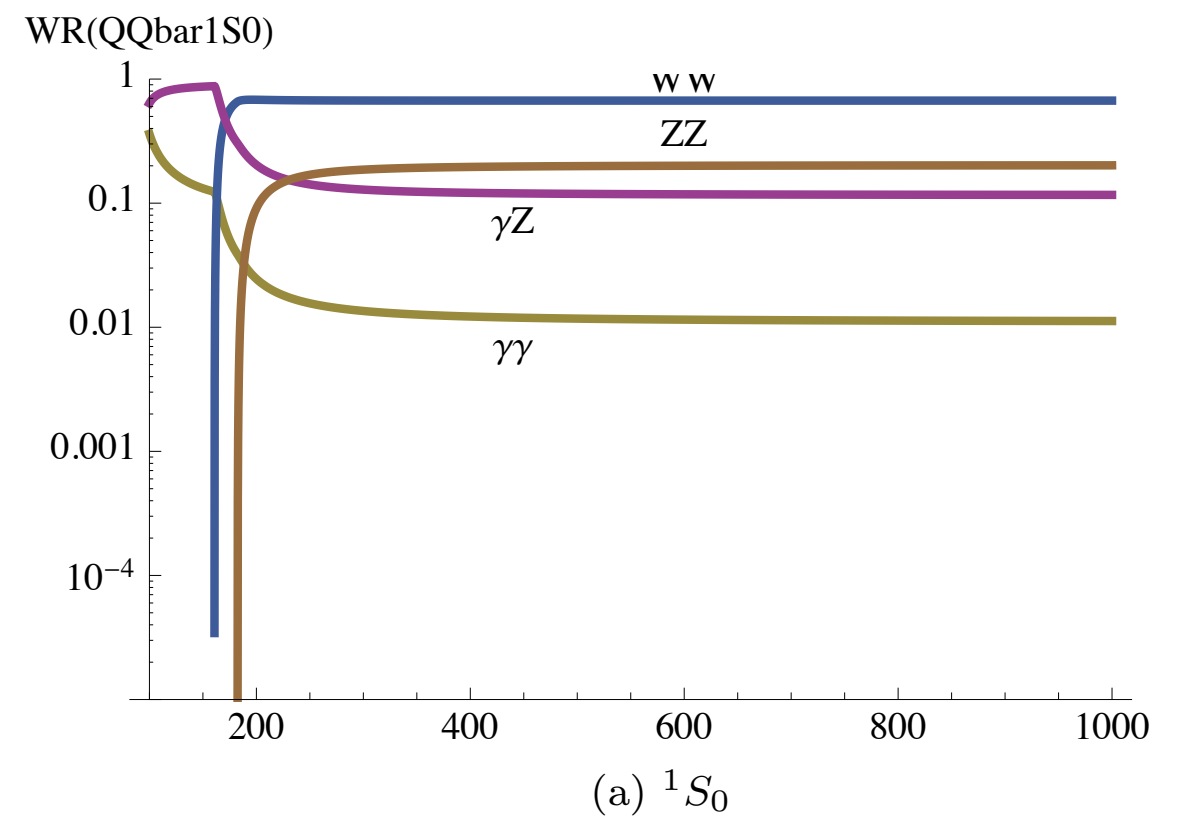
Missing energy signals largely absent!

Meson Decay Rates - A First Look

(Quirky) charged pion decay



(Vector-like) neutral meson decay



Fok, Kribs; 1106.3101

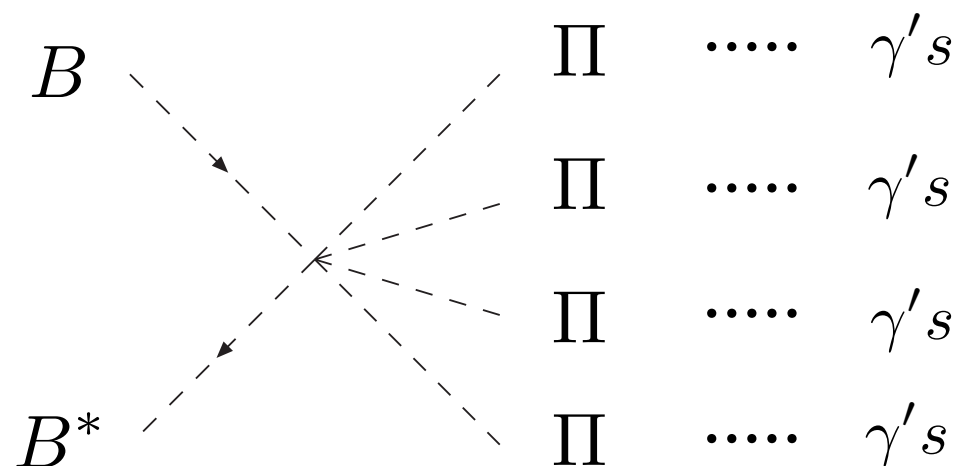
Astrophysical Signals - A First Look

Excited states of dark baryon that are nearby in mass

- fine structure
- hyperfine structure

could be visible through γ -ray emission/absorption lines.

If some symmetric component, annihilation signals (into γ s) are extremely interesting. It could be that multibody final states are generic, e.g.



Cascade annihilation begun
to be explored!

Elor, Rodd, Slatyer; 1503.01773

Summary and Future

- **Stealth Dark Matter** is a viable composite dark matter candidate composed of electrically charged constituents
- **Qualitatively different** direct, indirect, and collider signals are expected, illustrating the importance of “thinking outside the box”
- Meson production and decay is an extremely interesting LHC signal
 - calculating meson form factor f_{Π} from lattice is a high priority needed to make quantitative predictions for LHC
- S parameter from lattice would lead to bounds on EW breaking parameters (important!)
- Indirect astrophysical signals (γ -rays) possible between excited states as well as annihilation of a symmetric component
- Further investigation of abundance remains interesting (esp. asymmetric)